

# Geodesics

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- Straightest curves - longest proper times (locally)

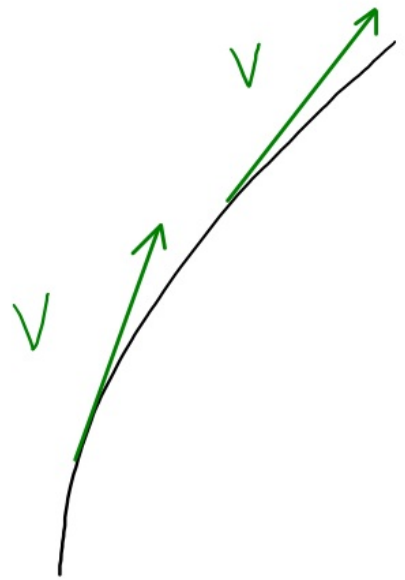
# Geodesics

- The paths of the free ... (to fall)
- Straightest curves - longest proper times (locally)
- Curvature makes parallel geodesics deviate  
\_ relative accelerations, "gravity"



\* a curve is a geodesic if it parallel transports its tangent vector:

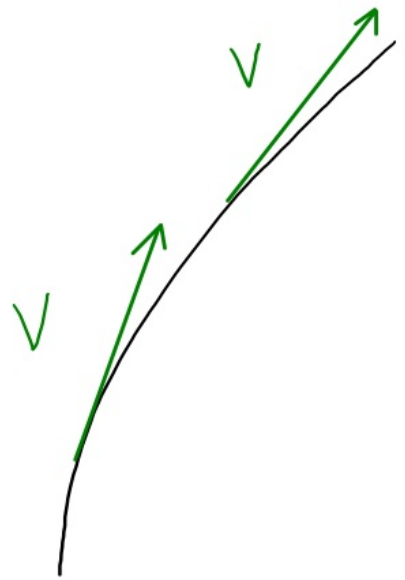
$$D_{\dot{\gamma}} \dot{\gamma}^{\mu} = 0 \Leftrightarrow \dot{\gamma}^{\nu} \nabla_{\nu} \dot{\gamma}^{\mu} = 0$$



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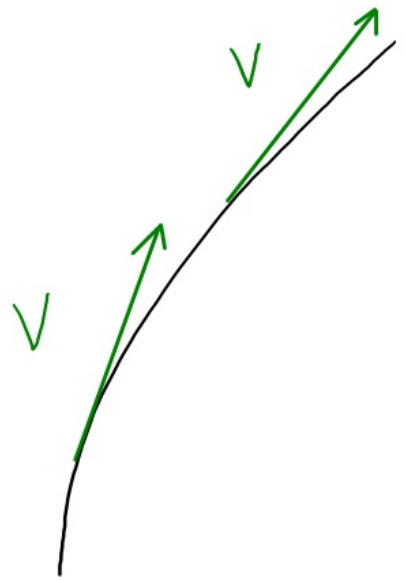
$$D_V V^\mu = 0 \Leftrightarrow V^\nu \nabla_\nu V^\mu = 0 \quad (1)$$

- a weaker condition is  $D_V V^\mu = f V^\mu$ , but with a reparametrization of the curve it can be recast to (1)



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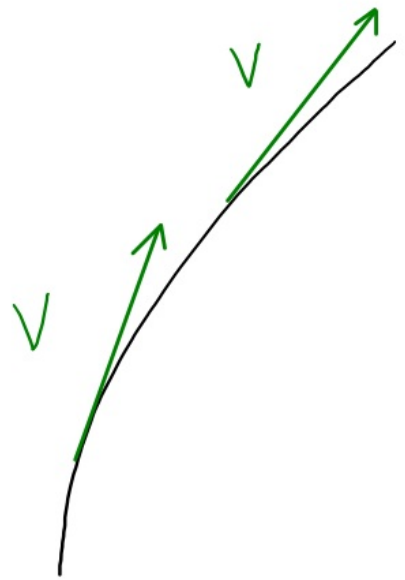


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- the parameter  $\tau$  in  $D_{\mathbf{v}} V^{\mu} = \frac{D V^{\mu}}{d\tau} = 0$  is an affine parameter

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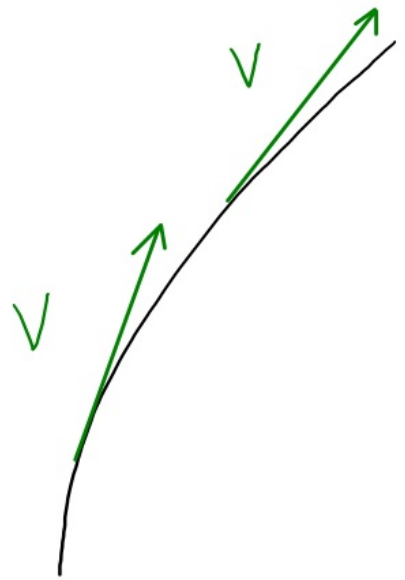
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- the parameter  $\tau$  in  $D_{\mathbf{v}} V^{\mu} = \frac{D V^{\mu}}{d\tau} = 0$  is an affine parameter  
 $\tau' = \alpha \tau + \beta$ ,  $\alpha, \beta \in \mathbb{R}$  is also an affine parameter

\* a curve is a geodesic if it parallel transports its tangent vector

$$D_{\dot{V}} V^{\mu} = 0 \Leftrightarrow V^{\nu} \nabla_{\nu} V^{\mu} = 0$$

$$\Rightarrow V^{\nu} \partial_{\nu} V^{\mu} + \Gamma^{\mu}_{\nu\rho} V^{\nu} V^{\rho} = 0$$





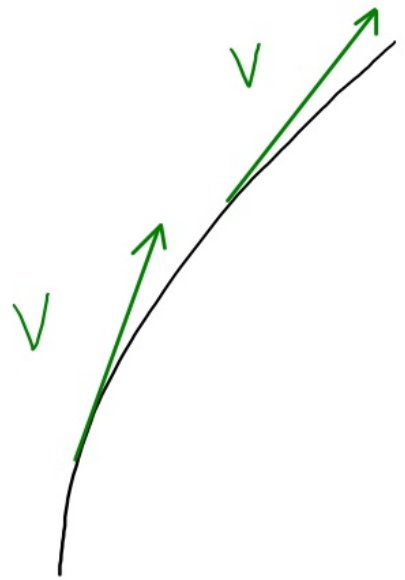
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If  $\{x^{\mu}\}$  are coordinates,  $V^{\mu} = \frac{dx^{\mu}}{d\tau}$ , and

$$\frac{dx^{\nu}}{d\tau} \frac{\partial}{\partial x^{\nu}} \left( \frac{dx^{\mu}}{d\tau} \right) + \Gamma^{\mu}_{\nu\rho} \frac{dx^{\nu}}{d\tau} \frac{dx^{\rho}}{d\tau} = 0$$



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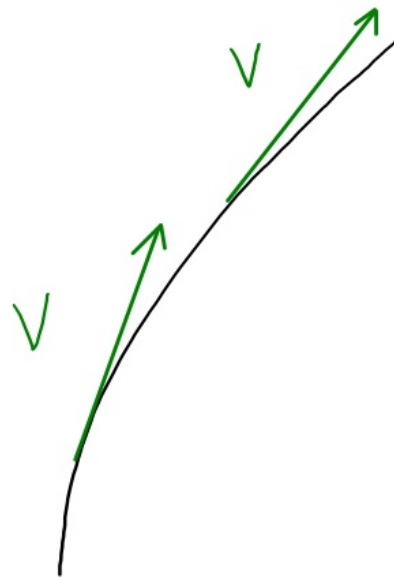
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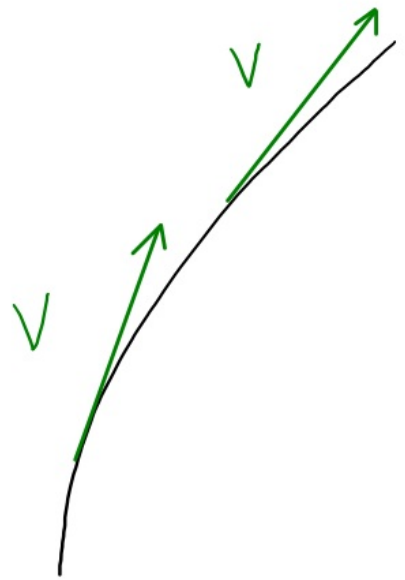
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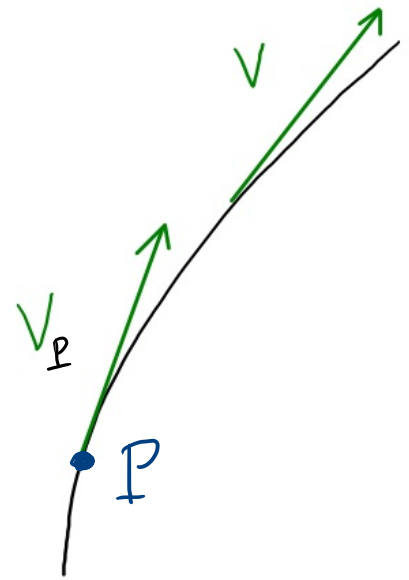
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$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma^\mu_{\nu\rho} \frac{dx^\nu}{d\tau} \frac{dx^\rho}{d\tau} = 0 \Rightarrow \left( x^\mu(\tau), \frac{dx^\mu}{d\tau}(\tau) \right) \rightarrow \text{unique solution}$$



$\Rightarrow$  There is a unique geodesic through  $P$  with tangent vector  $V_P^m$



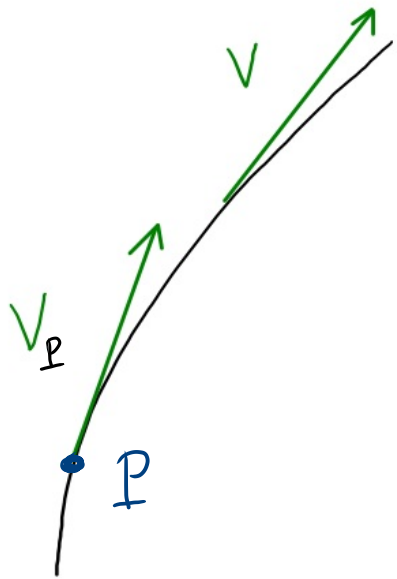
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$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma^\mu_{\nu\rho} \frac{dx^\nu}{d\tau} \frac{dx^\rho}{d\tau} = 0 \Rightarrow (x^\mu(0), \frac{dx^\mu}{d\tau}(0)) \rightarrow \text{unique solution}$$

⇒ There is a unique geodesic through  $P$  with tangent vector  $V_P^m$

⇒ Geodesics depend only on symmetric part of  $\Gamma^m_{\nu\rho}$

Adding  $\Gamma^m_{[\nu\rho]} \neq 0$  would not affect geodesics (torsion)



symmetric term under  $\nu \leftrightarrow \rho$

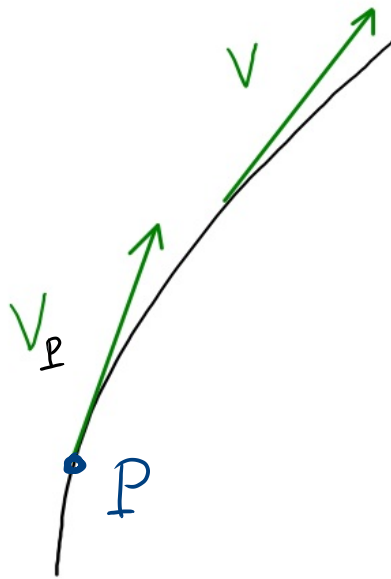
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⇒  $\Gamma^m_{\nu\rho} = 0$  everywhere  $\leadsto$  straight lines



↳ Flat spacetime has straight lines as geodesics!

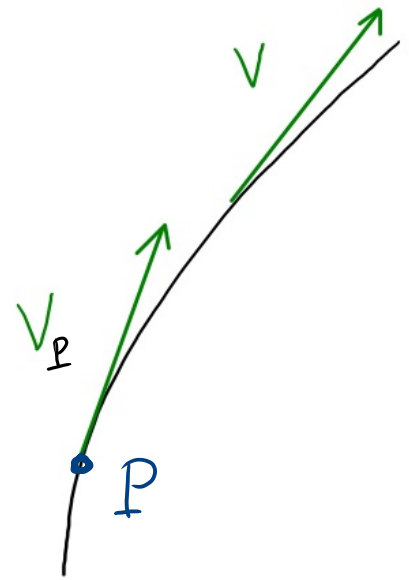
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↳ we can set  $\Gamma^m_{\nu\rho} = 0$  at a point  
"almost" straight lines!

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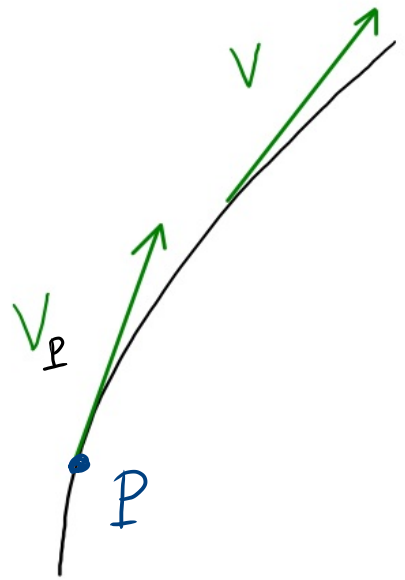
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• Free particles in GR move along geodesics

↳ Equivalence principle: Free particles appear to move on straight lines in inertial frames



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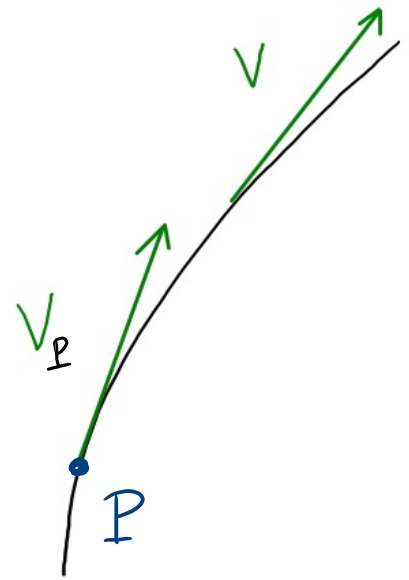
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- Free particles in GR move along geodesics

- Character of geodesics (timelike/null/spacelike) does not change

Inner product  $g_{\mu\nu} V^{\mu} V^{\nu}$  is constant:  $V^{\mu}$  parallel transported

• Free massive particles:

$$p^\mu = m U^\mu$$

$$U^\mu = \frac{dx^\mu}{dz}$$

4-velocity

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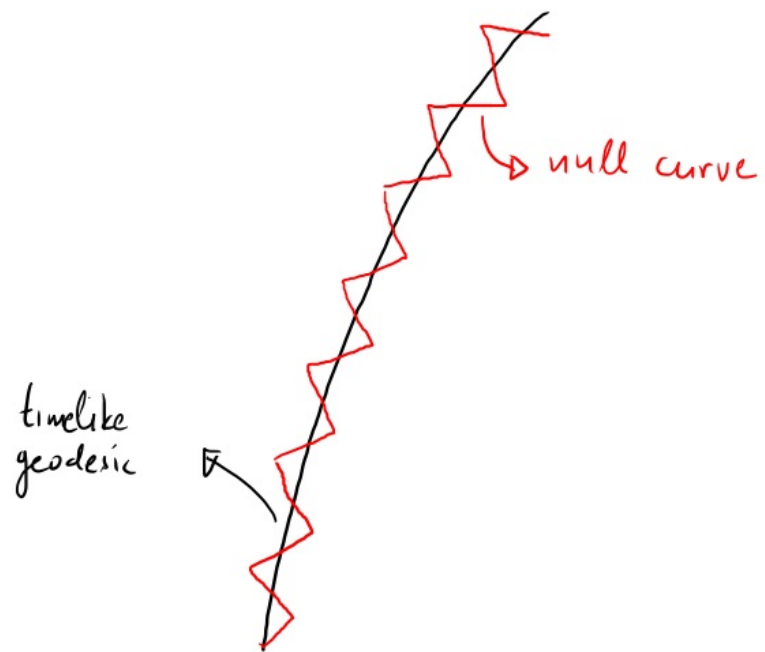
• Free massless particles:

We choose affine parameter  $\lambda$ , s.t.  $p^\mu = \frac{dx^\mu}{d\lambda}$

$$p^\nu \nabla_\nu p^\mu = 0$$

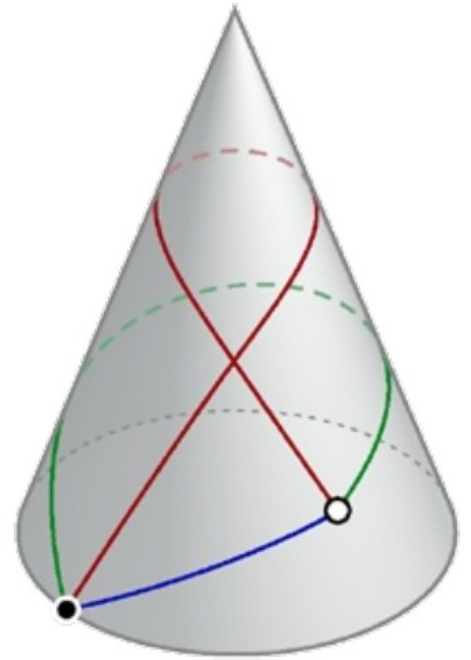
• Timelike geodesics are local maxima of proper time

(can't be minima, always  
arbitrarily close to a null curve  
- zero length - )



- Timelike geodesics are local maxima of proper time

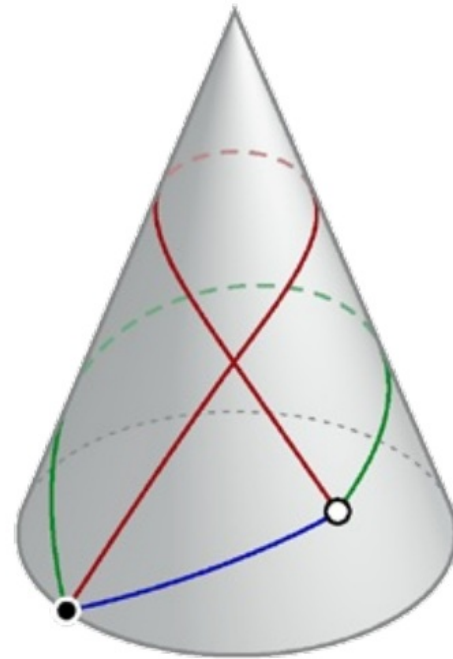
- Global topology may allow connecting two points with more than one geodesic
  - of different length -



<http://www.rdrop.com/~half/Creations/Puzzles/cone.geodesics/index.html>

• Timelike geodesics are local maxima  
of proper time  
hence the "local"

• Global topology may allow  
connecting two points with more  
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Extremization of length/proper time  $\rightarrow$  geodesics

Consider a timelike curve,  $ds^2 = -d\tau^2 < 0$

$$\tau = \int d\tau \left\{ -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}^{1/2}$$



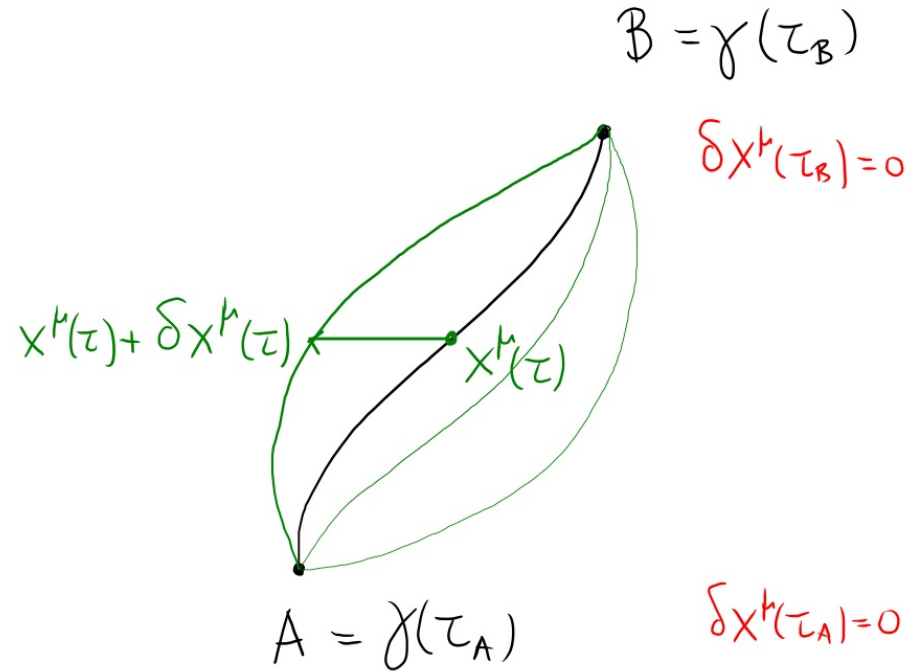
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Vary  $\gamma(\tau)$  s.t.  $\gamma'(\tau)$  is set of points with  $x^\mu(\tau) + \delta x^\mu(\tau)$ . Then

$$\delta\tau = \int d\tau \delta \left\{ -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}^{1/2}$$



# Extremization of length/proper time $\rightarrow$ geodesics

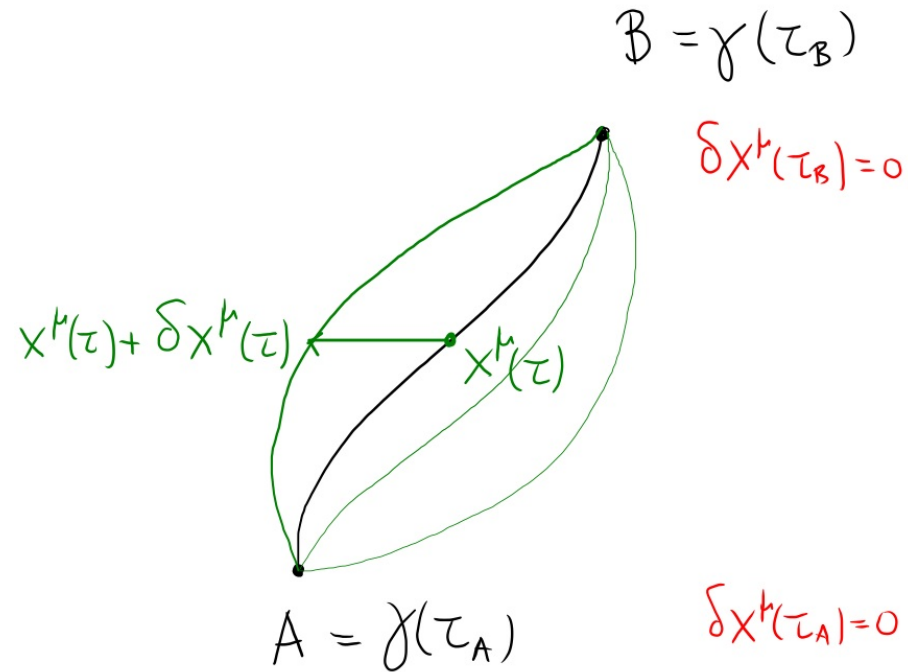
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$$\delta\tau = \int d\tau \delta \left\{ -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}^{1/2}$$

$$= -\frac{1}{2} \int d\tau \left\{ -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}^{-1/2} \delta \left\{ g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}$$



# Extremization of length/proper time $\rightarrow$ geodesics

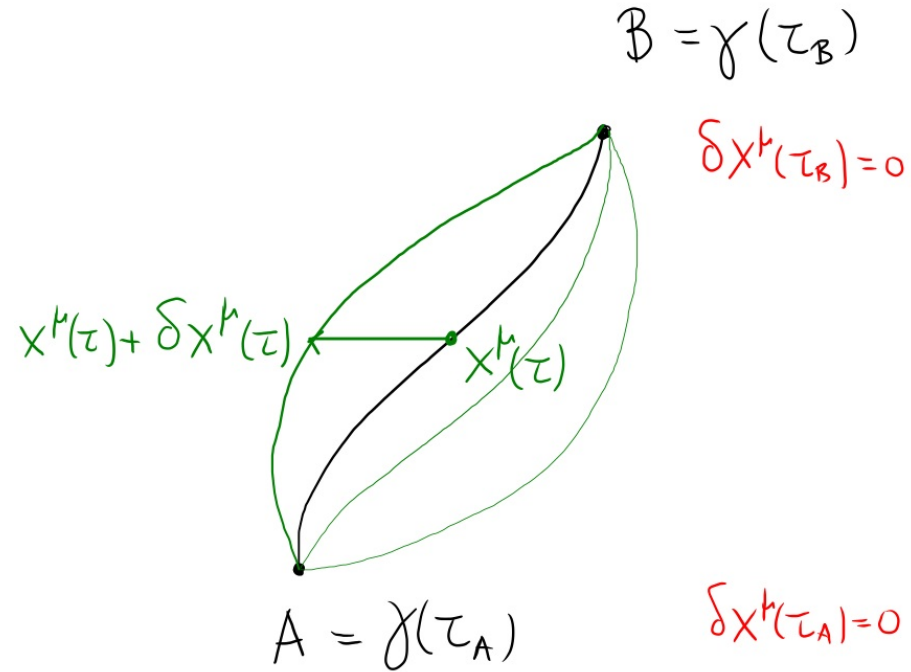
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Vary  $\gamma(\tau)$  s.t.  $\gamma'(\tau)$  is set of points with  $x^\mu(\tau) + \delta x^\mu(\tau)$ . Then

$$\begin{aligned} \delta\tau &= \int d\tau \delta \left\{ -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}^{1/2} \\ &= -\frac{1}{2} \int d\tau \delta \left\{ g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\}, \end{aligned}$$

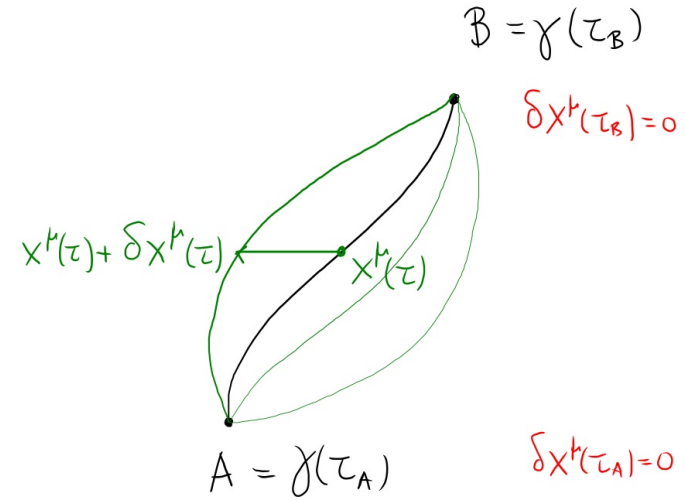
since  $g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} = g_{\mu\nu} U^\mu U^\nu = -1$



# Extremization of length/proper time $\rightarrow$ geodesics

$$x^\mu \rightarrow x^\mu + \delta x^\mu$$

$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$



$$\delta \tau = -\frac{1}{2} \int d\tau \delta \left\{ g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right\},$$

$$\text{since } g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} = g_{\mu\nu} U^\mu U^\nu = -1$$

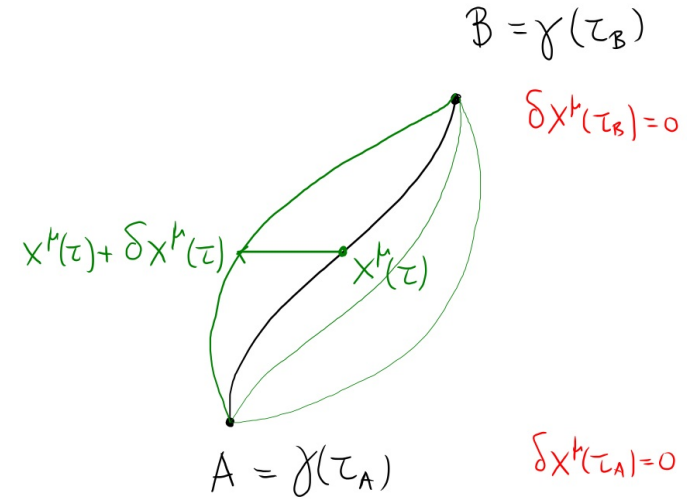
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$$\int d\tau \delta \left( g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) =$$

$$= \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\}$$



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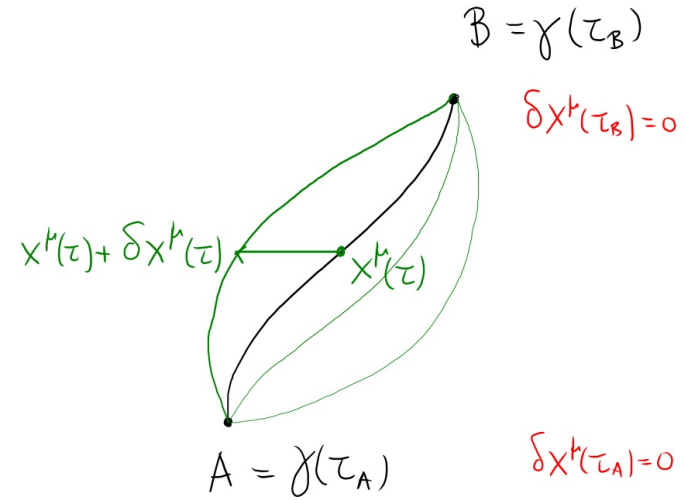
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- symmetric under  $\mu \leftrightarrow \nu$

$$\bullet \delta \left( \frac{dx^\mu}{d\tau} \right) = \frac{d}{d\tau} (\delta x^\mu)$$

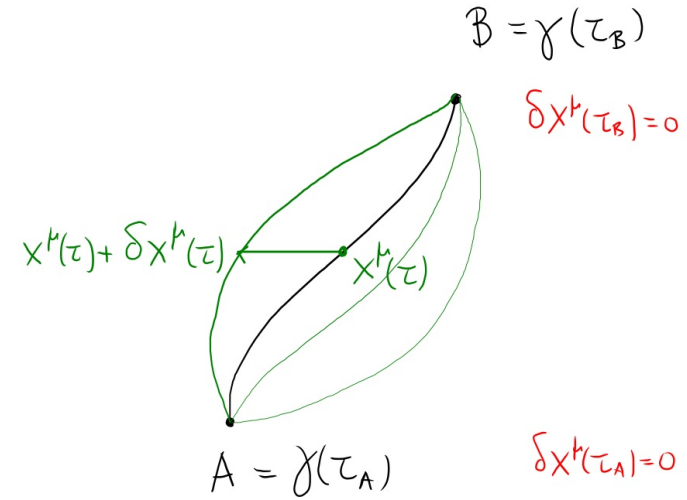


# Extremization of length/proper time $\rightarrow$ geodesics

$$x^\mu \rightarrow x^\mu + \delta x^\mu$$

$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

$$\begin{aligned} & \int d\tau \delta \left( g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) = \\ & = \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\} \\ & = \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau} \right\} \end{aligned}$$



# Extremization of length/proper time $\rightarrow$ geodesics

$$X^\mu \rightarrow X^\mu + \delta X^\mu$$

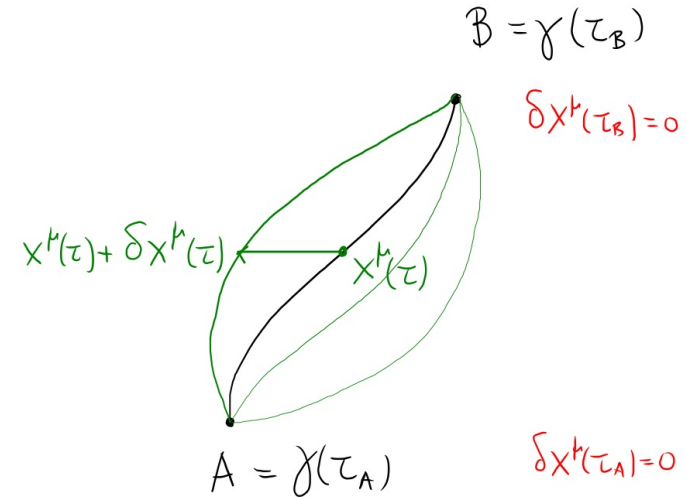
$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

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$$\text{But } 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d\delta x^\nu}{d\tau} = \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta x^\nu \right] - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu$$





# Extremization of length/proper time $\rightarrow$ geodesics

$$X^\mu \rightarrow X^\mu + \delta X^\mu$$

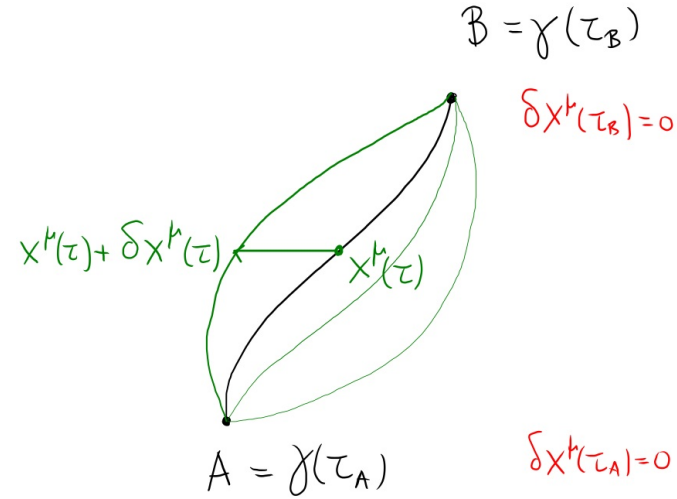
$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

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$$\int_{\tau_A}^{\tau_B} 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d\delta X^\nu}{d\tau} = \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta X^\nu \right] \Big|_{\tau_A}^{\tau_B} - \int_{\tau_A}^{\tau_B} \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta X^\nu d\tau$$



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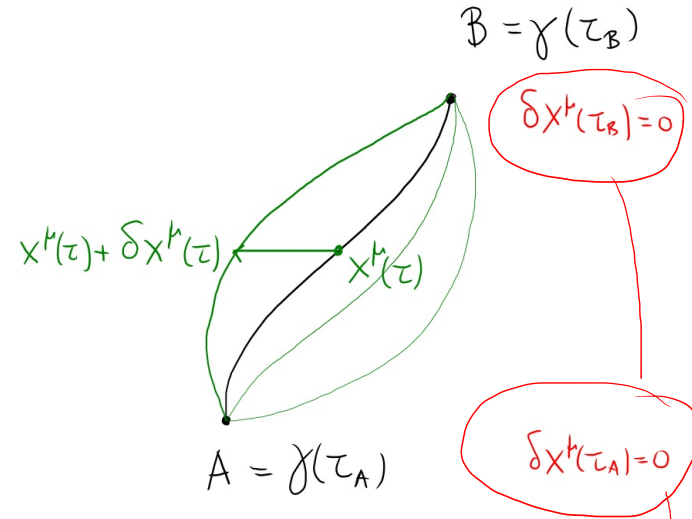
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$$\int d\tau \delta \left( g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) =$$

$$= \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\}$$

$$= \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d}{d\tau} (\delta x^\nu) \right\}$$

$$\int_{\tau_A}^{\tau_B} 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d}{d\tau} \delta x^\nu = \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta x^\nu \right]_{\tau_A}^{\tau_B} - \int_{\tau_A}^{\tau_B} \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu d\tau$$



# Extremization of length/proper time $\rightarrow$ geodesics

$$x^\mu \rightarrow x^\mu + \delta x^\mu$$

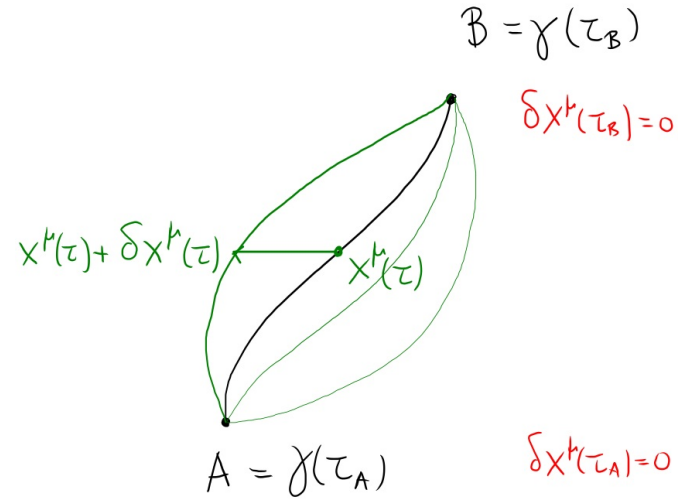
$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

$$\int d\tau \delta \left( g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) =$$

$$= \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\}$$

$$= \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau} \right\}$$

$$= \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$



# Extremization of length/proper time $\rightarrow$ geodesics

$$x^\mu \rightarrow x^\mu + \delta x^\mu$$

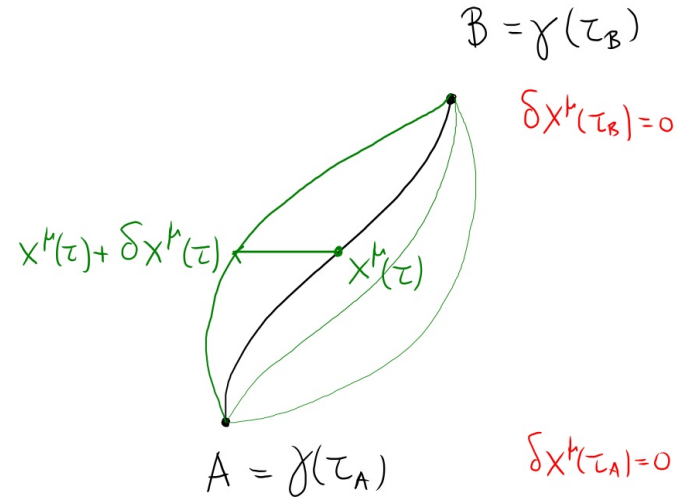
$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

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$$= \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\}$$

$$= \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau} \right\}$$

$$= \int d\tau \left\{ \partial^\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$



# Extremization of length/proper time $\rightarrow$ geodesics

$$x^\mu \rightarrow x^\mu + \delta x^\mu$$

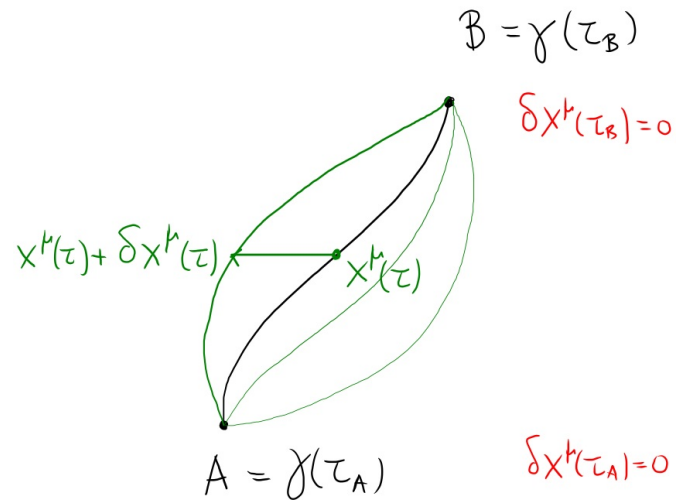
$$g_{\mu\nu} \rightarrow g_{\mu\nu} + \partial_\lambda g_{\mu\nu} \delta x^\lambda + \dots \quad (\text{Taylor series expansion})$$

$$\int d\tau \delta \left( g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) =$$

$$= \int d\tau \left\{ \delta g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + g_{\mu\nu} \delta \left( \frac{dx^\mu}{d\tau} \right) \frac{dx^\nu}{d\tau} + g_{\mu\nu} \frac{dx^\mu}{d\tau} \delta \left( \frac{dx^\nu}{d\tau} \right) \right\}$$

$$= \int d\tau \left\{ \partial_\lambda g_{\mu\nu} \delta x^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{d(\delta x^\nu)}{d\tau} \right\}$$

$$= \int d\tau \left\{ \partial_\nu g_{\mu\lambda} \delta x^\nu \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$



Extremization of length/proper time  $\rightarrow$  geodesics

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$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \left[ \frac{d}{d\tau} g_{\mu\nu} \right] \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

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$$\delta\tau = \int d\tau \left\{ \partial_\nu g_{\mu\lambda} \delta x^\nu \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$

Extremization of length/proper time  $\rightarrow$  geodesics

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$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \left[ \frac{d}{d\tau} g_{\mu\nu} \right] \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

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$$\delta\tau = \int d\tau \left\{ \partial_\nu g_{\mu\lambda} \delta x^\nu \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$

Extremization of length/proper time  $\rightarrow$  geodesics

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \left[ \frac{d}{d\tau} g_{\mu\nu} \right] \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \partial_\lambda g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

symmetric under  $\mu \leftrightarrow \lambda$

$$= - \int d\tau \delta x^\nu \left\{ 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} + (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\mu\lambda}) \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} \right\}$$

$\mu \leftrightarrow \lambda$

$$\delta \tau = \int d\tau \left\{ \partial_\nu g_{\mu\lambda} \delta x^\nu \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - \frac{d}{d\tau} \left[ 2 g_{\mu\nu} \frac{dx^\mu}{d\tau} \right] \delta x^\nu \right\}$$



Extremization of length/proper time  $\rightarrow$  geodesics

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \left[ \frac{d}{d\tau} g_{\mu\nu} \right] \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \partial_\lambda g_{\mu\nu} \frac{dx^\lambda}{d\tau} \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

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Extremization  $\delta\tau = 0$  for any  $\delta x^\mu$  implies the integrand is zero:

$$2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} + (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\mu\lambda}) \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} = 0$$

Extremization of length/proper time  $\rightarrow$  geodesics

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \left[ \frac{d}{d\tau} g_{\mu\nu} \right] \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

$$= \int d\tau \delta x^\nu \left\{ \partial_\nu g_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} - 2 \partial_\lambda g_{\mu\nu} \frac{dx^\lambda}{d\tau} \frac{dx^\mu}{d\tau} - 2 g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} \right\}$$

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$$g^{\sigma\mu} g_{\mu\nu} \frac{d^2 x^\mu}{d\tau^2} + \frac{1}{2} g^{\sigma\mu} (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\mu\lambda}) \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} = 0$$

Extremization of length/proper time  $\rightarrow$  geodesics

$$\frac{d^2 x^\sigma}{d\tau^2} + \Gamma^\sigma_{\mu\lambda} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} = 0$$

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$$\underbrace{g^{\sigma\mu} g_{\mu\nu}}_{\delta^\sigma_\nu} \frac{d^2 x^\mu}{d\tau^2} + \frac{1}{2} \underbrace{g^{\sigma\mu} (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\mu\lambda})}_{\Gamma^\sigma_{\mu\lambda}} \frac{dx^\mu}{d\tau} \frac{dx^\lambda}{d\tau} = 0$$

Example: Flat space cosmology

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$$ds^2 = -dt^2 + a^2(t) [dx^2 + dy^2 + dz^2]$$

Example: Flat space cosmology

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$$ds^2 = -dt^2 + a^2(t) [dx^2 + dy^2 + dz^2]$$

$$= -dt^2 + a^2(t) \delta_{ij} dx^i dx^j \quad i, j = 1, 2, 3$$

Example: Flat space cosmology

$$ds^2 = -dt^2 + a^2(t) [dx^2 + dy^2 + dz^2]$$

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Calculate  $\Gamma^{\lambda}_{\mu\nu}$ :

$$I = \frac{1}{2} \int d\tau \left[ - \left( \frac{dt}{d\tau} \right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$x^\mu \rightarrow x^\mu + \delta x^\mu, \quad \delta I = 0$$

Example: Flat space cosmology

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$$(a) \quad t \rightarrow t + \delta t \quad \Rightarrow \quad \delta a(t) = \frac{da}{dt} \delta t = \dot{a}(t) \delta t$$

Example: Flat space cosmology

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$$\delta I = \frac{1}{2} \int d\tau \left[ -2 \frac{dt}{d\tau} \frac{d\delta\tau}{d\tau} + \delta a^2 \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$



Example: Flat space cosmology

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$$= \frac{1}{2} \int d\tau \left[ +2 \frac{d^2 t}{d\tau^2} \delta t + 2a \dot{a} \delta t \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

---

(a)  $t \rightarrow t + \delta t \quad \Rightarrow \quad \delta a(t) = \frac{da}{dt} \delta t = \dot{a}(t) \delta t$

$$\delta I = \frac{1}{2} \int d\tau \left[ -2 \frac{dt}{d\tau} \frac{d\delta\tau}{d\tau} + \delta a^2 \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

Example: Flat space cosmology

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$$= \frac{1}{2} \int d\tau \left[ 2 \frac{d^2 t}{d\tau^2} \delta t + 2a \dot{a} \delta t \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \left[ \frac{d^2 t}{d\tau^2} + a \dot{a} \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \delta t$$

---

(a)  $t \rightarrow t + \delta t \quad \Rightarrow \quad \delta a(t) = \frac{da}{dt} \delta t = \dot{a}(t) \delta t$

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Example: Flat space cosmology

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$$= \frac{1}{2} \int d\tau \left[ 2 \frac{d^2 t}{d\tau^2} \delta t + 2a \dot{a} \delta t \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \left[ \frac{d^2 t}{d\tau^2} + a \dot{a} \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \delta t$$

$$\delta I = 0 \quad \forall \delta t \Rightarrow \frac{d^2 t}{d\tau^2} + a \dot{a} \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} = 0$$

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Example: Flat space cosmology

$$= \frac{1}{2} \int d\tau \left[ 2 \frac{d^2 t}{d\tau^2} \delta t + 2a \dot{a} \delta t \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \left[ \frac{d^2 t}{d\tau^2} + a \dot{a} \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \delta t$$

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compare with

$$\frac{d^2 t}{d\tau^2} + \Gamma^0_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} = 0$$

Example: Flat space cosmology

$$= \frac{1}{2} \int d\tau \left[ 2 \frac{d^2 t}{d\tau^2} \delta t + 2a \dot{a} \delta t \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \left[ \frac{d^2 t}{d\tau^2} + a \dot{a} \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \delta t$$

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compare with

$$\frac{d^2 t}{d\tau^2} + \Gamma^0_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} = 0$$

$$\Rightarrow \Gamma^0_{00} = \Gamma^0_{0i} = \Gamma^0_{i0} = 0$$
$$\Gamma^0_{ij} = a \dot{a} \delta_{ij}$$

Example: Flat space cosmology

$$I = \frac{1}{2} \int d\tau \left[ -\left(\frac{dt}{d\tau}\right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

$$x^i \rightarrow x^i + \delta x^i$$

Example: Flat space cosmology

$$I = \frac{1}{2} \int d\tau \left[ -\left(\frac{dt}{d\tau}\right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \quad x^i \rightarrow x^i + \delta x^i$$

$$\delta I = \frac{1}{2} \int d\tau \left[ 0 + a^2(t) \delta_{ij} \delta \left[ \frac{dx^i}{d\tau} \right] \frac{dx^j}{d\tau} + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right] \right]$$

# Example: Flat space cosmology

$$I = \frac{1}{2} \int d\tau \left[ -\left(\frac{dt}{d\tau}\right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right]$$

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$$= \frac{1}{2} \int d\tau \ 2 a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right]$$



Example: Flat space cosmology

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# Example: Flat space cosmology

$$I = \frac{1}{2} \int d\tau \left[ -\left(\frac{dt}{d\tau}\right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \quad x^i \rightarrow x^i + \delta x^i$$

$$\delta I = \frac{1}{2} \int d\tau \left[ 0 + a^2(t) \delta_{ij} \delta \left[ \frac{dx^i}{d\tau} \right] \frac{dx^j}{d\tau} + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right] \right]$$

$$= \frac{1}{2} \int d\tau \ 2 a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \ a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{d}{d\tau} \delta x^j$$

$$\begin{aligned} & \delta x^j(\tau_A) = \delta x^j(\tau_B) = 0 \\ & \int_{\tau_A}^{\tau_B} a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta x^j \end{aligned}$$

$$= - \int d\tau \ \frac{d}{d\tau} \left[ a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \right] \delta x^j + \int d\tau \ \frac{d}{d\tau} \left[ a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta x^j \right]$$

# Example: Flat space cosmology

$$I = \frac{1}{2} \int d\tau \left[ -\left(\frac{dt}{d\tau}\right)^2 + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{dx^j}{d\tau} \right] \quad x^i \rightarrow x^i + \delta x^i$$

$$\delta I = \frac{1}{2} \int d\tau \left[ 0 + a^2(t) \delta_{ij} \delta \left[ \frac{dx^i}{d\tau} \right] \frac{dx^j}{d\tau} + a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right] \right]$$

$$= \frac{1}{2} \int d\tau \ 2 a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \delta \left[ \frac{dx^j}{d\tau} \right]$$

$$= \int d\tau \ a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \frac{d}{d\tau} \delta x^j$$

$$= - \int d\tau \ \frac{d}{d\tau} \left[ a^2(t) \delta_{ij} \frac{dx^i}{d\tau} \right] \delta x^j$$

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## Example: Flat space cosmology

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Example: Flat space cosmology

---

$$\Rightarrow \Gamma^i_{00} = \Gamma^i_{jk} = 0$$

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Null geodesics: consider photons moving on +X axis

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solution:  $\frac{dt}{d\lambda} = \frac{w_0}{a(t)}$

• consider comoving observers:  $U^\mu = (1, 0, 0, 0)$

$$E_{\text{photon}} = -p_\mu U^\mu = -g_{00} p^0 U^0 = -g_{00} \frac{dx^0}{d\tau} \cdot 1 = + \frac{\omega_0}{a}$$

solution:  $\frac{dt}{d\tau} = \frac{\omega_0}{a(t)}$

•  $a(t)$  is known from Einstein equations, (3)  $\Rightarrow t=t(\tau)$  and (1)  $\Rightarrow x=x(\tau)$



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$$\frac{\omega_2}{\omega_1} = \frac{\omega_0/a_2}{\omega_0/a_1} = \frac{a_1}{a_2} \quad \text{cosmological redshift}$$

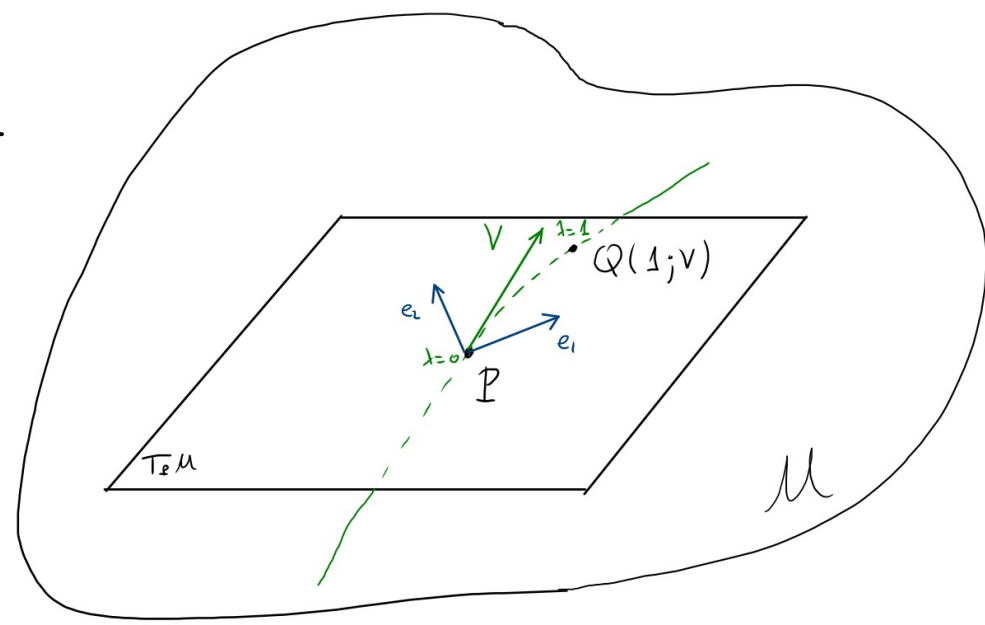
Note:

- it is the energy of the photon that is reduced
- photon is not "stretched"
- the relation of  $E$  and  $\omega$  is quantum mechanical

# Riemann Normal Coordinates

Construct local inertial frame using geodesics:

- Pick an event  $P$  and consider all geodesics through  $P$



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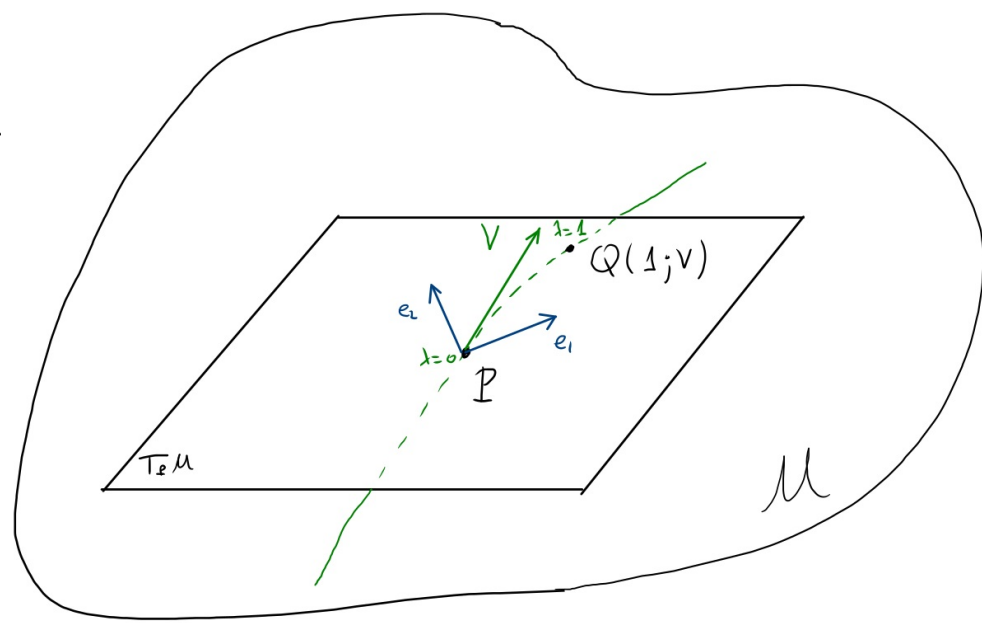
- Pick an event  $P$  and consider all geodesics through  $P$

- Each vector  $V$  at  $P$  determines a geodesic w/ affine parameter  $\lambda$

Consider the point  $Q = Q(\lambda; V)$

where on geodesic  $\leftarrow$

$\leftarrow$  which geodesic



# Riemann Normal Coordinates

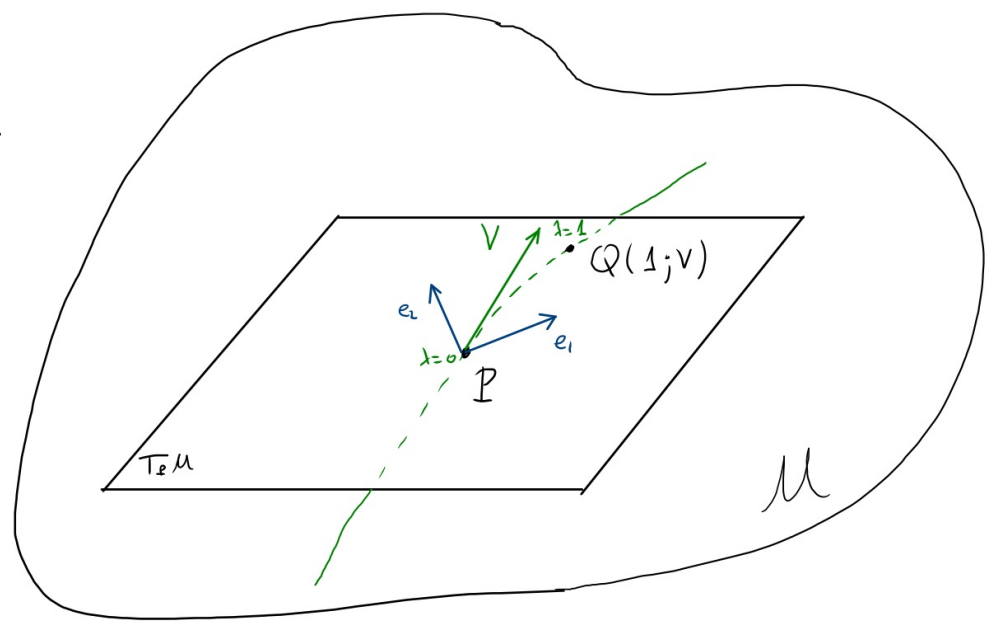
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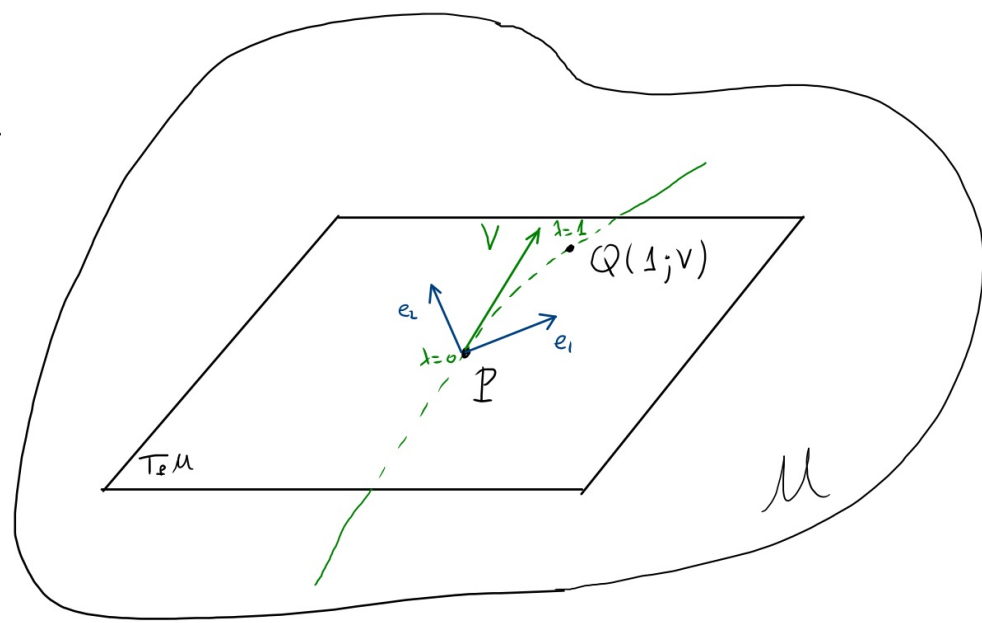
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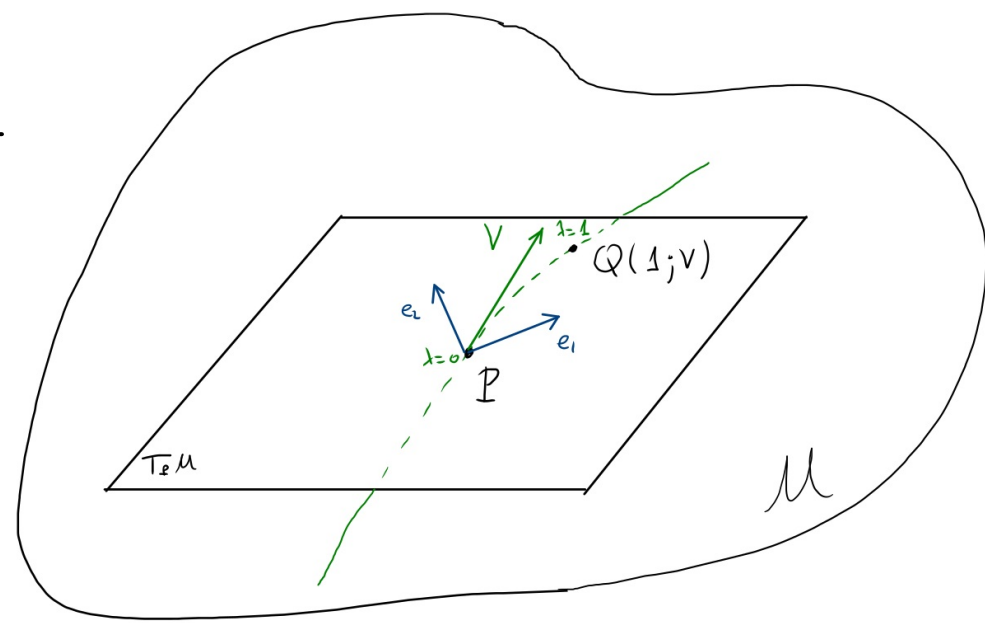
- As long as geodesics don't intersect, we have 1-1 map of points + vectors



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⇒ Choose orthonormal basis  $\{e_\mu\}$  at  $T_P \mathcal{M}$

If  $V = x^\mu e_\mu$ , define coordinates of  $Q(1; V)$  to be  $\{x^\mu\}$



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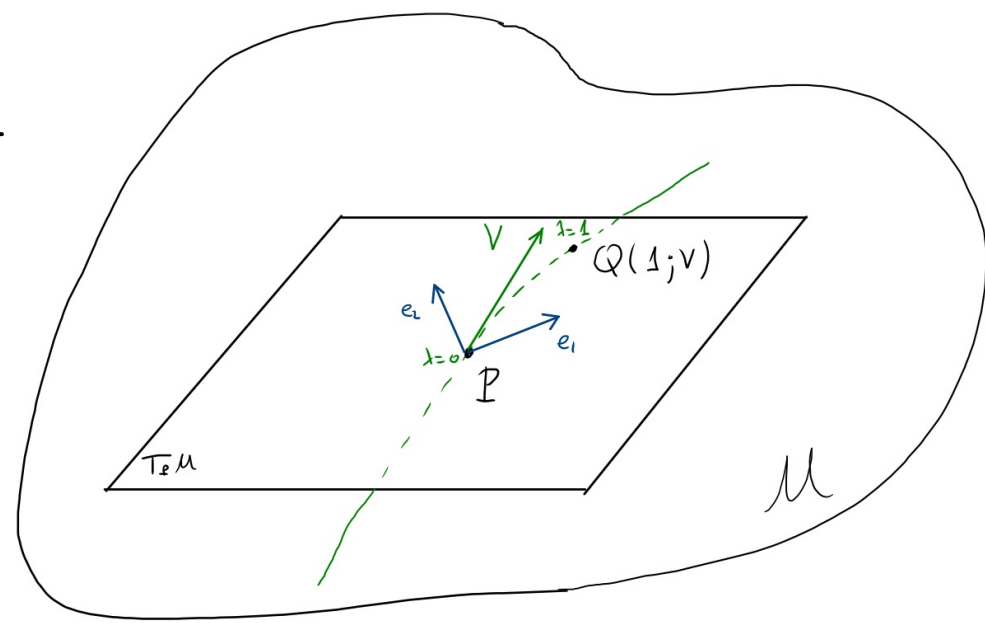
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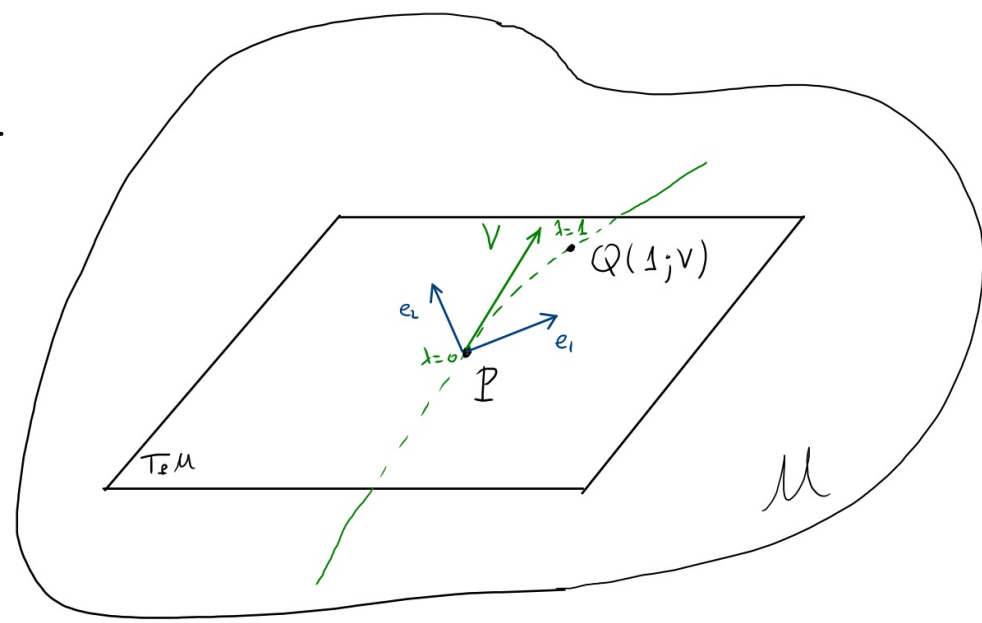
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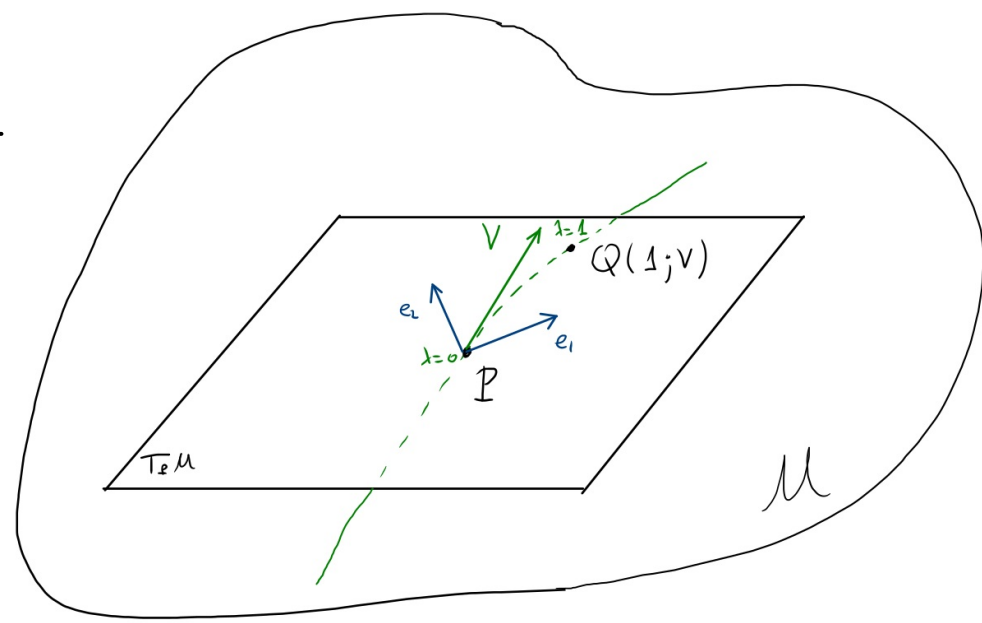
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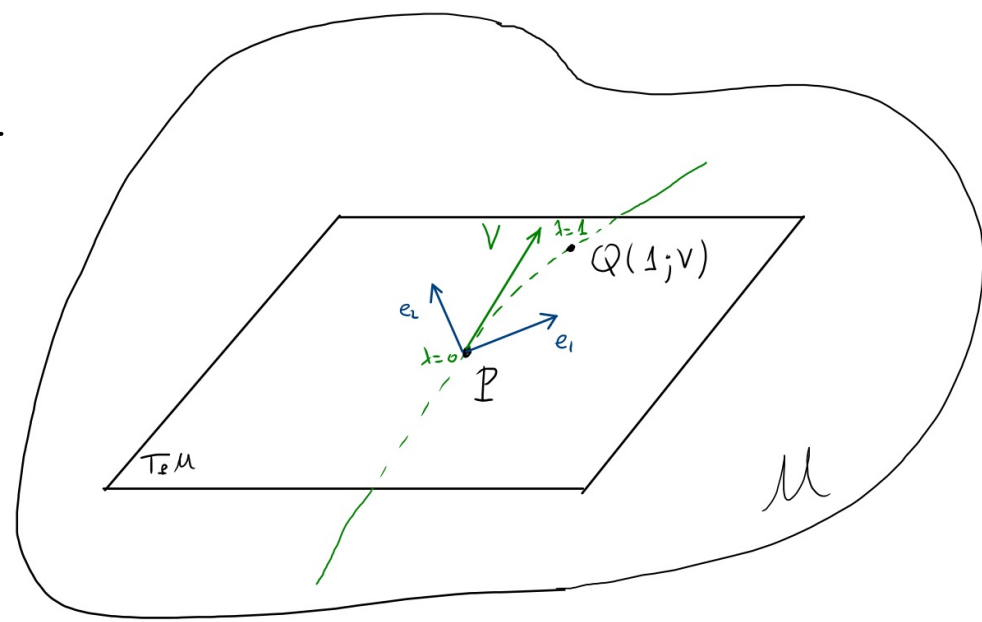
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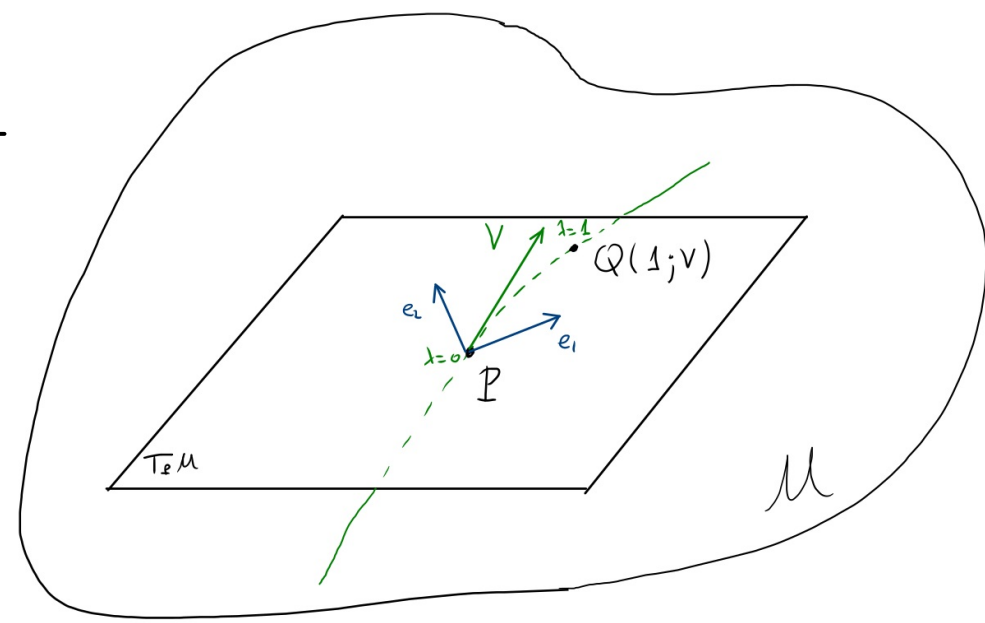
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Higher derivatives determined by curvature:

$$\partial_\rho \partial_\sigma g_{\mu\nu}(P) = -\frac{1}{3} (R_{\mu\sigma\nu\rho} + R_{\mu\rho\nu\sigma})$$

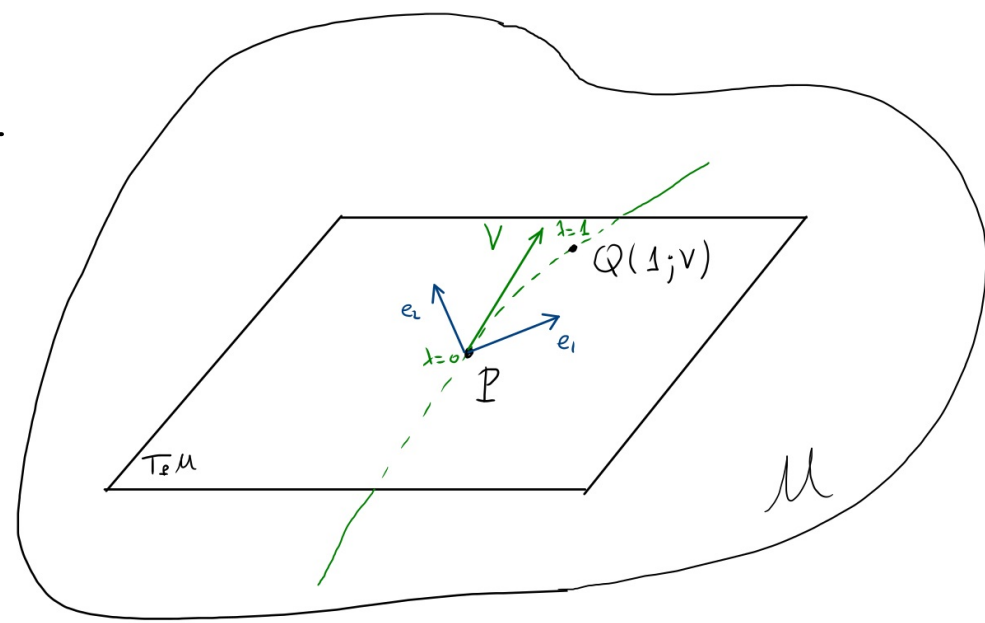
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• Any other normal coordinate system same to 2<sup>nd</sup> order

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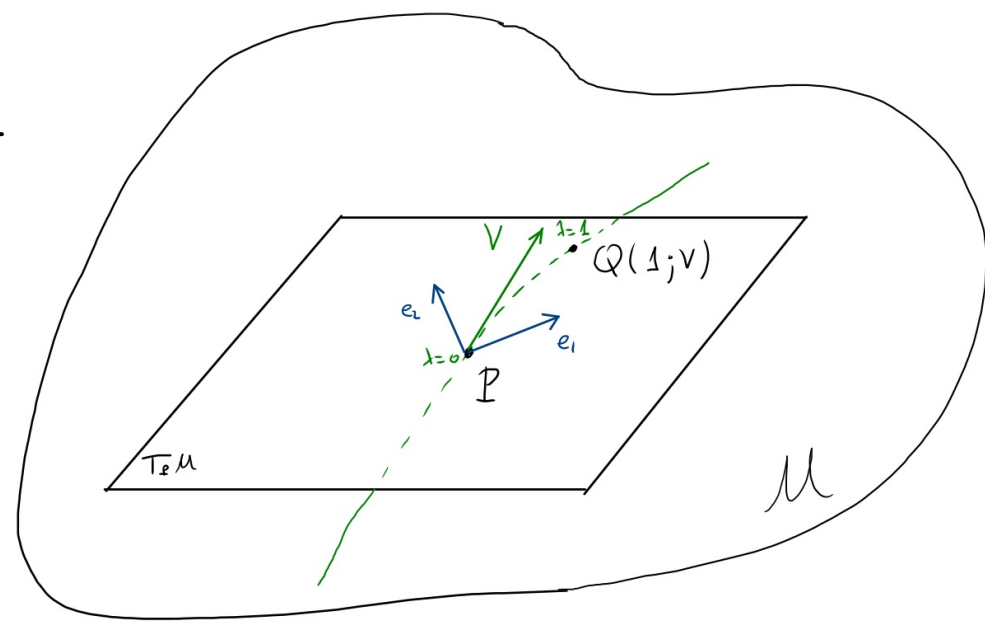
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• Any other normal coordinate system same to 2<sup>nd</sup> order

$$x^\mu(\rho) = x^\mu(P) + \mathcal{O}(\rho^3)$$

• If same to 3<sup>rd</sup> order,  
Eqs (1) + (2) are preserved

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Example: Wormhole geometry (Hartle, examples 8.2, 8.3, 8.5)

$$ds^2 = -dt^2 + dr^2 + (b^2 + r^2)(d\theta^2 + \sin^2\theta d\phi^2)$$

$$g_{tt} = -1 \quad g_{rr} = 1$$

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easier: only  $g_{tt}$ ,  $g_{rr}$ ,  $g_{\theta\theta}$ ,  $g_{\phi\phi}$

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$\mu\nu\rho$  + cyclic 

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$\mu\nu\rho$  + cyclic

only permutations  
of  $r\theta\theta, r\phi\phi, \theta\phi\phi$   
may appear

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

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$$g_{m\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\rho m} + \partial_\rho g_{m\nu} - \partial_m g_{\nu\rho})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$

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$$g_{m\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} ( \partial_\nu g_{\rho m} + \partial_\rho g_{\nu m} - \partial_m g_{\nu\rho} )$$

$$g \Gamma = \frac{1}{2} ( \partial g + \partial g - \partial g )$$

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$$g_{r\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\rho r} + \partial_\rho g_{r\nu} - \partial_r g_{\nu\rho})$$

$$g_r \Gamma_{\theta\theta} = \frac{1}{2} (\partial_\theta g_{\theta r} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta})$$

$$g_\theta \Gamma_{r\theta} = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta})$$

$$g_r \Gamma_{\phi\phi} = \frac{1}{2} (\partial_\phi g_{\phi r} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi})$$

$$g_\phi \Gamma_{r\phi} = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi})$$

$$g_\theta \Gamma_{\phi\phi} = \frac{1}{2} (\partial_\phi g_{\phi\theta} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi})$$

$$g_\phi \Gamma_{\theta\phi} = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi})$$

permute cyclically:  
 start from 3<sup>rd</sup> term,  
 then 2<sup>nd</sup>, then 1<sup>st</sup>

$r \rightarrow \theta \rightarrow \phi$   
 $\theta \rightarrow r \rightarrow \theta$   
 $r \rightarrow \phi \rightarrow \phi$   
 $\phi \rightarrow r \rightarrow \phi$   
 $\theta \rightarrow \phi \rightarrow \phi$   
 $\phi \rightarrow \theta \rightarrow \phi$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} ( \partial_\nu g_{\mu\rho} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho} )$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} ( \partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta} )$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} ( \partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta} )$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} ( \partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi} )$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} ( \partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi} )$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} ( \partial_\phi g_{\theta\phi} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi} )$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} ( \partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi} )$$

no other choice:  $g_{\mu\nu}$  is diagonal



$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\rho\mu} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta})$$

Non-zero terms

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta})$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi})$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\phi} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\mu\rho} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta})$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi})$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\theta} + \partial_\phi g_{\theta\theta} - \partial_\theta g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\mu\rho} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta}) \Rightarrow (b^2 + r^2) \Gamma_{r\theta}^\theta = \frac{1}{2} \cdot 2r \Rightarrow \Gamma_{r\theta}^\theta = \frac{r}{b^2 + r^2}$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi})$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\phi} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\mu\rho} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta}) \Rightarrow (b^2 + r^2) \Gamma_{r\theta}^\theta = \frac{1}{2} \cdot 2r \Rightarrow \Gamma_{r\theta}^\theta = \frac{r}{b^2 + r^2}$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi}) \Rightarrow 1 \cdot \Gamma_{\phi\phi}^r = \frac{1}{2} (-2r \sin^2 \theta) \Rightarrow \Gamma_{\phi\phi}^r = -r \sin^2 \theta$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi})$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\theta} + \partial_\phi g_{\theta\theta} - \partial_\theta g_{\phi\phi})$$

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$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\rho\mu} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{r\theta} - \partial_\theta g_{r\theta}) \Rightarrow (b^2 + r^2) \Gamma_{r\theta}^\theta = \frac{1}{2} \cdot 2r \Rightarrow \Gamma_{r\theta}^\theta = \frac{r}{b^2 + r^2}$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi}) \Rightarrow 1 \cdot \Gamma_{\phi\phi}^r = \frac{1}{2} (-2r \sin^2 \theta) \Rightarrow \Gamma_{\phi\phi}^r = -r \sin^2 \theta$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{r\phi} - \partial_\phi g_{r\phi}) \Rightarrow (b^2 + r^2) \sin^2 \theta \cdot \Gamma_{r\phi}^\phi = \frac{1}{2} 2r \sin^2 \theta \Rightarrow \Gamma_{r\phi}^\phi = \frac{r}{b^2 + r^2}$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\theta} + \partial_\phi g_{\theta\theta} - \partial_\theta g_{\phi\phi})$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\theta\phi} - \partial_\phi g_{\theta\phi})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\rho\mu} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{r\theta} - \partial_\theta g_{r\theta}) \Rightarrow (b^2 + r^2) \Gamma_{r\theta}^\theta = \frac{1}{2} \cdot 2r \Rightarrow \Gamma_{r\theta}^\theta = \frac{r}{b^2 + r^2}$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi}) \Rightarrow 1 \cdot \Gamma_{\phi\phi}^r = \frac{1}{2} (-2r \sin^2 \theta) \Rightarrow \Gamma_{\phi\phi}^r = -r \sin^2 \theta$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{r\phi} - \partial_\phi g_{r\phi}) \Rightarrow (b^2 + r^2) \sin^2 \theta \cdot \Gamma_{r\phi}^\phi = \frac{1}{2} 2r \sin^2 \theta \Rightarrow \Gamma_{r\phi}^\phi = \frac{r}{b^2 + r^2}$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\theta} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi}) \Rightarrow (b^2 + r^2) \Gamma_{\phi\phi}^\theta = \frac{1}{2} (-2(b^2 + r^2) \sin \theta \cos \theta) \Rightarrow \Gamma_{\phi\phi}^\theta = -\sin \theta \cos \theta$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\theta\phi} - \partial_\phi g_{\theta\phi})$$

$$g_{rr} = 1 \quad g_{\theta\theta} = b^2 + r^2 \quad g_{\phi\phi} = (b^2 + r^2) \sin^2 \theta$$

$$\partial_r g_{\theta\theta} = 2r \quad \partial_r g_{\phi\phi} = 2r \sin^2 \theta \quad \partial_\theta g_{\phi\phi} = 2(b^2 + r^2) \sin \theta \cos \theta$$


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$$g_{\mu\sigma} \Gamma_{\nu\rho}^\sigma = \frac{1}{2} (\partial_\nu g_{\mu\rho} + \partial_\rho g_{\mu\nu} - \partial_\mu g_{\nu\rho})$$

$$g_{rr} \Gamma_{\theta\theta}^r = \frac{1}{2} (\partial_\theta g_{r\theta} + \partial_\theta g_{r\theta} - \partial_r g_{\theta\theta}) \Rightarrow 1 \cdot \Gamma_{\theta\theta}^r = \frac{1}{2} (-2r) \Rightarrow \Gamma_{\theta\theta}^r = -r$$

$$g_{\theta\theta} \Gamma_{r\theta}^\theta = \frac{1}{2} (\partial_r g_{\theta\theta} + \partial_\theta g_{\theta r} - \partial_\theta g_{r\theta}) \Rightarrow (b^2 + r^2) \Gamma_{r\theta}^\theta = \frac{1}{2} \cdot 2r \Rightarrow \Gamma_{r\theta}^\theta = \frac{r}{b^2 + r^2}$$

$$g_{rr} \Gamma_{\phi\phi}^r = \frac{1}{2} (\partial_\phi g_{r\phi} + \partial_\phi g_{r\phi} - \partial_r g_{\phi\phi}) \Rightarrow 1 \cdot \Gamma_{\phi\phi}^r = \frac{1}{2} (-2r \sin^2 \theta) \Rightarrow \Gamma_{\phi\phi}^r = -r \sin^2 \theta$$

$$g_{\phi\phi} \Gamma_{r\phi}^\phi = \frac{1}{2} (\partial_r g_{\phi\phi} + \partial_\phi g_{\phi r} - \partial_\phi g_{r\phi}) \Rightarrow (b^2 + r^2) \sin^2 \theta \cdot \Gamma_{r\phi}^\phi = \frac{1}{2} 2r \sin^2 \theta \Rightarrow \Gamma_{r\phi}^\phi = \frac{r}{b^2 + r^2}$$

$$g_{\theta\theta} \Gamma_{\phi\phi}^\theta = \frac{1}{2} (\partial_\phi g_{\theta\theta} + \partial_\phi g_{\theta\phi} - \partial_\theta g_{\phi\phi}) \Rightarrow (b^2 + r^2) \Gamma_{\phi\phi}^\theta = \frac{1}{2} (-2(b^2 + r^2) \sin \theta \cos \theta) \Rightarrow \Gamma_{\phi\phi}^\theta = -\sin \theta \cos \theta$$

$$g_{\phi\phi} \Gamma_{\theta\phi}^\phi = \frac{1}{2} (\partial_\theta g_{\phi\phi} + \partial_\phi g_{\phi\theta} - \partial_\phi g_{\theta\phi}) \Rightarrow (b^2 + r^2) \sin^2 \theta \Gamma_{\theta\phi}^\phi = \frac{1}{2} 2(b^2 + r^2) \sin \theta \cos \theta \Rightarrow \Gamma_{\theta\phi}^\phi = \cot \theta$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2 + r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2 + r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2 + r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2 + r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2 + r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2 + r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

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$$\ddot{r} + \Gamma^r_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2+r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2+r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

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$$\ddot{r} + \Gamma^r_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow$$

$$\ddot{r} + \Gamma^r_{\theta\theta} \dot{\theta} \dot{\theta} + \Gamma^r_{\phi\phi} \dot{\phi} \dot{\phi} = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2 + r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2 + r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

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$$\ddot{r} + \Gamma^r_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow$$

$$\ddot{r} + \Gamma^r_{\theta\theta} \dot{\theta}^2 + \Gamma^r_{\phi\phi} \dot{\phi}^2 = 0 \Rightarrow$$

$$\ddot{r} - r \dot{\theta}^2 - r \sin^2\theta \dot{\phi}^2 = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2+r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2+r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

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$$\ddot{r} + \Gamma^r_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow$$

$$\ddot{r} + \Gamma^r_{\theta\theta} \dot{\theta} \dot{\theta} + \Gamma^r_{\phi\phi} \dot{\phi} \dot{\phi} = 0 \Rightarrow$$

$$\ddot{r} - r \dot{\theta}^2 - r \sin^2\theta \dot{\phi}^2 = 0$$

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$$\ddot{\theta} + \Gamma^\theta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

$$\Gamma^\theta_{r\theta} = \Gamma^\theta_{\theta r} = \frac{r}{b^2 + r^2} \quad \Gamma^\theta_{\phi\phi} = -\sin\theta \cos\theta$$

$$\Gamma^\phi_{r\phi} = \Gamma^\phi_{\phi r} = \frac{r}{b^2 + r^2} \quad \Gamma^\phi_{\phi\theta} = \Gamma^\phi_{\theta\phi} = \cot\theta$$

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$$\ddot{t} + \Gamma^t_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow \ddot{t} = 0$$

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$$\ddot{r} + \Gamma^r_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow$$

$$\ddot{r} + \Gamma^r_{\theta\theta} \dot{\theta} \dot{\theta} + \Gamma^r_{\phi\phi} \dot{\phi} \dot{\phi} = 0 \Rightarrow$$

$$\ddot{r} - r \dot{\theta}^2 - r \sin^2\theta \dot{\phi}^2 = 0$$

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$$\ddot{\theta} + \Gamma^\theta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu = 0 \Rightarrow$$

$$\ddot{\theta} + \Gamma^\theta_{r\theta} \dot{r} \dot{\theta} + \Gamma^\theta_{\theta r} \dot{\theta} \dot{r} + \Gamma^\theta_{\phi\phi} \dot{\phi} \dot{\phi} = 0$$

$$\Gamma^r_{\theta\theta} = -r \quad \Gamma^r_{\phi\phi} = -r \sin^2\theta$$

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Radially falling particle through wormhole:

$$\ddot{\theta} = \dot{\phi} = 0$$

$$(1) \Rightarrow t = \alpha \tau + \beta \quad (\text{proper time too})$$

$$(2) \Rightarrow \ddot{r} = 0$$

$$\ddot{t} = 0 \quad (1)$$

$$\ddot{r} - r \dot{\theta}^2 - r \sin^2 \theta \dot{\phi}^2 = 0 \quad (2)$$

$$\ddot{\theta} + \frac{2r}{b^2 + r^2} \dot{r} \dot{\theta} - \sin \theta \cos \theta \dot{\phi}^2 = 0$$

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Consider 4-velocity ( $u > 0$ )

$$u^\mu = (u^t, u^r, 0, 0)$$

$$\ddot{t} = 0 \quad (1)$$

$$\ddot{r} - r \dot{\theta}^2 - r \sin^2 \theta \dot{\phi}^2 = 0 \quad (2)$$

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Consider 4-velocity ( $u > 0$ )

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$$-(u^0)^2 + (u^r)^2 = -1 \Rightarrow$$

$$u^0 = \sqrt{1 + (u^r)^2}$$

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Radially falling particle through wormhole:

$$\ddot{\theta} = \dot{\phi} = 0$$

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Consider 4-velocity ( $u > 0$ )

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take  $u^\mu = (\sqrt{1+U^2}, U, 0, 0)$ , then

$$(3) \Rightarrow r(\tau) = U \cdot \tau$$

if  $r(0) = 0$  (i.e. passing the throat)

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So, going from  $r(\tau_1) = R$  to

$r(\tau_2) = -R$ , the proper time for the particle is

$$\Delta \tau = \tau_2 - \tau_1 = \frac{2R}{U}$$

Radially falling particle through wormhole:

The spatial distance it travelled was ( $dt = d\theta = d\phi = 0$ )

$$s = \int ds = \int |dr| = 2R$$

take  $u^\mu = (\sqrt{1+U^2}, U, 0, 0)$ , then

$$(3) \Rightarrow r(\tau) = U \cdot \tau$$

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